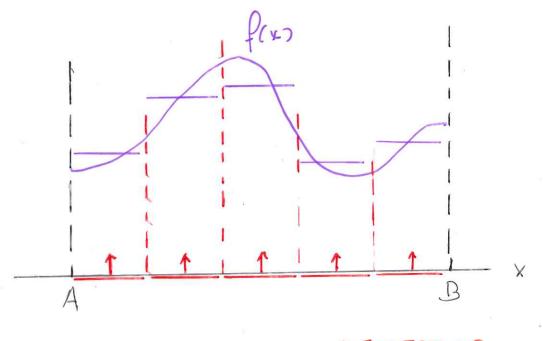
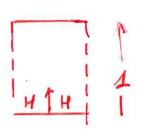
INVERSIONES NUMERICAS

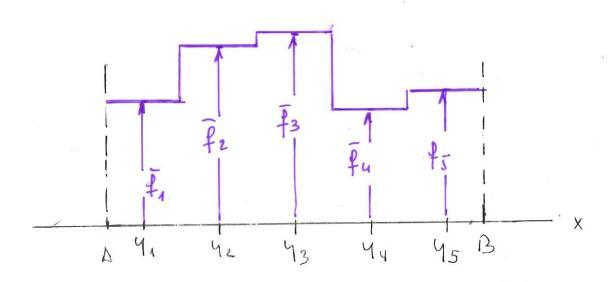
13.1

INVERSION NUMERICA "NAIF"





DETECTOR IDEAL, CON GSA FORMA



$$y_{1} = A + H$$

$$y_{2} = y_{1} + 2H$$

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$$y_{4} = y_{3} + 2H$$

$$y_{7} = y_{4} + 2H$$

$$y_{7} + H = B$$

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$$f_{2} = g(y_{2}) = \int_{y_{1} - H}^{y_{2} + H} f(x) dx = 2H$$

$$f_{3} = g(y_{2}) = \int_{y_{3} - H}^{y_{4} + H} f(x) dx = 2H$$

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 $Y_{3} = Y_{2} + \delta = A - H + 3\delta = \Delta + \delta$
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$$9_{5} = 5(7_{5}) = \int_{16-10}^{7_{5+10}} f(x) dx = -8_{1} + 5_{4} + 5_{5} + 5_{5}$$

$$7_{6+10} = 5(7_{10} + 5_{10} + 5_{10} + 5_{10} + 5_{10} + 5_{10} + 5_{10} + 5_{10}$$

$$9_{6} = 5(7_{10}) = \int_{16-10}^{7_{6+10}} f(x) dx = -(8_{10} - 9_{10}) + 8_{10} + 5_{10} + 5_{10}$$

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INVERSA

$$\begin{aligned}
F_{1} &= \frac{1}{5}S_{1} \\
F_{2} &= \frac{1}{5}(S_{1}-S_{1}) \\
F_{3} &= \frac{1}{5}(S_{3}-S_{2}) \\
F_{4} &= \frac{1}{5}(S_{4}-S_{3}) \\
F_{5} &= \frac{1}{5}(S_{5}-S_{4}+S_{4}) \\
F_{6} &= \frac{1}{5}(S_{6}-S_{7}+S_{2}-S_{4}) \\
F_{7} &= \frac{1}{5}(S_{7}-S_{6}+S_{3}-S_{2}) \\
F_{8} &= \frac{1}{5}(S_{7}-S_{7}+S_{4}-S_{3})
\end{aligned}$$

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13.2

INVERSION MATRICIAL

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NUCLEO VIA INTEGRACIÓN

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ADMITIENDO UNA FORMULA DE INTEGRACIÓN NUMERICA PARA LA INTEGRAL SOBRE

TOMANDO TANTOS VALORES DE Y Y DE G(Y)

COMO SE HAN ELEGIDO PARA LA INTEGRAL

SOBRE X : EN TERMINOS DE ORDENADAS

DISCRETAS X (UN NUMERO MAYOR PARA

Y LUDIERA SER MÁS CONVENIENTE AUN

- MINIMOS CUADRADOS - PERO CUIDADO NO

REFETIR ECUACIONES)

TENDREMOS UN SISTEMA LINEAL
DE CUYA INVERSION SE DEDUCIRÁ
EL CONJUNTO DE VALORES $f(x_J)$

Inverse Problems in Astronomy

a guide to inversion strategies for remotely sensed data

Ian J D Craig

BSc, PhD

Senior Lecturer in Mathematics University of Waikato, Hamilton, New Zealand

John C Brown
BSc, PhD, DSc, FRAS, FRSE

Professor of Astrophysics University of Glasgow

Adam Hilger Ltd, Bristol and Boston

5.2 Numerical Inversion Methods

To illustrate the basic methods involved it is sufficient to consider the reduction of the Fredholm integral equation of the first kind

$$\int_{a}^{b} k(x, y) f(y) \, \mathrm{d}y = g(x) \qquad a \leqslant x \leqslant b \tag{5.1}$$

to matrix form. Equations of the Volterra type can generally be accommodated within the same basic framework (see e.g. §5.4).

It is instructive, first of all, to comment on the effect of specifying the data function only at discrete points, x_i say, where $a \le x_1 < \ldots < x_i < \ldots < x_m \le b$. Equation (5.1) then becomes

$$\int_{a}^{b} k(x_{i}, y) f(y) dy = g(x_{i}) \qquad i = 1, 2, ..., m.$$
 (5.2)

Although no error is involved in this discretisation it is clear that many functions f(y) can satisfy equation (5.2) without being solutions of equation (5.1). This reflects the fact that the discretised data points can never uniquely represent the data function, in that, any polynomial of degree (m-1) or greater can pass through the node points of g. Thus a strong element of information loss has already crept into the functional problem by its discretisation over the x variable.

It follows that the practical inversion problem can never possess an unambiguous solution, even in the absence of data noise.

5.2.1 Matrix-quadrature methods

This is probably the simplest of all numerical methods. The integral in equation (5.2) is approximated directly by a quadrature sum so that

$$\sum_{j=1}^{n} k(x_i, y_j) w_j f(y_j) \simeq g(x_i) \qquad i = 1, 2, ..., m$$
 (5.3)

where the w_j are weighting coefficients associated with the quadrature rule. A first-order approximation is obtained in the simplest case of uniform weighting $(w_j = \Delta y = (b-a)/(n-1))$ but more accuracy can, in principle, be derived by employing the trapezoidal rule or the Simpson quadrature (see Chapter 6 for detailed application). These correspond respectively, to piecewise constant, linear and quadratic approximations to the integrand in equation (5.2). Still greater accuracy in representation (5.3) can be derived by employing more sophisticated quadrature formulae, for example those derived by Gaussian integration techniques. In all cases, however, an $m \times n$ matrix system is derived, i.e.

$$\sum_{j=1}^{n} A_{ij} f_j = g_i \qquad i = 1, 2, ..., m$$
 (5.4)

pes 103

The central problem is then to determine the differential emission measure function $\xi(T)$, describing the distribution of material over temperature in the source, from the bremsstrahlung spectrum $\varphi(\varepsilon_i)$. The 'integral' equation to be inverted (cf equation (2.32)) takes the discrete form

$$\varphi(\varepsilon_i) = \int_T k(\varepsilon_i, T) \, \xi(T) \, dT \qquad (7.9)$$

for each ε_i at which φ is measured, with the kernel

$$k(\varepsilon, T) = \frac{A}{T^{1/2}} \exp(-\varepsilon/T)$$
 (7.10)

where A is a scaling factor and T is in energy units.

In practice it is often convenient to express equation (7.9) as a standard Laplace transform (cf equation (2.33))

$$\varphi(\varepsilon_i) = \int_0^\infty \exp(-\varepsilon_i t) \ G(t) \ \mathrm{d}t. \tag{7.11}$$

7.3.2 Classical inversions

per 196

(a) Matrix-quadrature. We begin by exploring the following problem. Suppose we are given n noise free measurements of the power law spectrum equation (7.13) in some typical energy range, say 3-10 keV. Can we obtain a satisfactory n-point approximation to the true solution, equation (7.14) by invoking the classical quadrature technique? (In what follows we discretise equation (7.9) using the Simpson quadrature, generally taking constant logarithmic intervals in the $\varepsilon - T$ variables (see §7.4.2).)

First we observe from figure 7.3 that the linear system obtained by discretising equation (7.9) is generally extremely badly conditioned, even for relatively low order numerical approximations (n = 5-9 say). Curve B for example, illustrates the behaviour of the condition number C_A for the energy band 3–10 keV and the temperature span 5–60 × 10⁶ K. Of special interest is the rapid (exponential) rise in C_A as the number of quadrature points is increased. We see by equation (5.27) that to guarantee a stable solution in only a nine point approximation requires relative data errors approaching 1 part in 10⁸, obviously a quite impossible requirement. Nor is the matter radically improved by choosing different wavelength and temperature bands (curves A and C): the rapid rise in C_A as the discretisation mesh is 'filled in' is a universal trend that is not much affected by the temperature—photon energy representation.

that the data function namely

$$\varphi(s) = \int_0^\infty \exp(-st)N(t) dt \qquad (7.40)$$

is given by the Laplace transform of the source function. In constructing numerical solutions we deal only with discretised approximations to equation (7.40).

7.4.2 Numerical inversion method

We employ two methods to reduce equation (7.40) to matrix form

$$\varphi_i = \sum_{j=1}^n K_{ij} N_j \qquad i = 1, \dots, m.$$
 (7.41)

First we invoke the first-order product integration method introduced in $\S5.2.3$ (see $\S\S5.3.3$ (Example A) and 5.4.3 for applications). Although only exact for piecewise constant source functions, this method has the attraction of not requiring a uniformly incremented t mesh, nor any strategem for truncating the infinite t integration interval. The kernel matrix is given by

$$K_{ij} = \frac{1}{s_i} \left[\exp(-s_i t_{j-1}) - \exp(-s_i t_j) \right]$$
 (7.42)

where an individual partition $(t_j - t_{j-1})$ can be chosen in *any* convenient manner.

Of course greater accuracy can be derived by employing a conventional matrix quadrature, for example Simpson's rule. In this case the fundamental t interval must be partitioned using equal increments with a consequent loss of flexibility in the representation of the source function. Moreover, in practice, the large variation over both frequency (ν) and energy (γ) makes it expedient to introduce logarithmic variables (as in §7.3.1)

$$x = \ln(s) \qquad \qquad y = \ln(t) \tag{7.43}$$

so that the kernel in equation (7.40) becomes

$$k(x, y) = \exp(y - e^{x+y})$$
 (7.44)

where $-\infty < y < \infty$. Assuming an appropriate truncation of this interval (see below) the kernel is discretised by taking a uniform increment Δy whence

$$K_{ij} = w_j k(x_i, y_i) \Delta y \tag{7.45}$$

 w_j being the weights associated with Simpson's rule namely (1, 2, 4, 2, ..., 2, 4, 2, 1) with n odd. The inversion problem is then reduced to equation (7.41) on identifying $\varphi_i \equiv \exp x_i$ and $N_j \equiv N \exp y_j$. This system is formally far more accurate than equation (7.42), being exact for cubic rather than zeroth-order polynomial source functions.

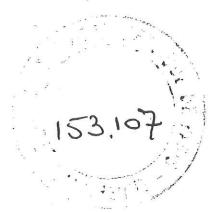
Two strategies will be adopted to stabilise the matrix inversion problem.

THE NUMERICAL TREATMENT OF INTEGRAL EQUATIONS

By

CHRISTOPHER T. H. BAKER M.A., D.Phil.

Reader in Mathematics, University of Manchester



CLARENDON PRESS · OXFORD 1977

5.16. Quadrature methods and expansion methods

Let us investigate briefly how a quadrature method may be set up for an equation of the first kind. We shall then indicate the *potentially* unsatisfactory nature of the method. Modifications of the method, intended to alleviate some of its shortcomings, will be discussed later.

From eqn (5.104) we have, on replacing the integral by a rule with weights w_j and abscissae y_j $(j=0,1,2,\ldots,n)$, the approximation

$$\sum_{j=0}^{n} w_{j} K(x, y_{j}) f(y_{j}) \simeq g(x) .$$

If we write

$$\sum_{j=0}^{n} w_{j} K(x, y_{j}) \tilde{f}(y_{j}) = g(x)$$
 (5.107)

and seek values of $\tilde{f}(y_j)$ $(j=0,1,2,\ldots,n)$ satisfying (5.107), we shall not in general find a solution. A solution exists only if g(x) is a linear combination of the functions $K(x,y_j)$ $(j=0,1,2,\ldots,n)$ and, even if the integral equation has a solution, this is not guaranteed. However, if we select z_0, z_1, \ldots, z_m we may be able to satisfy the equations

$$\sum_{j=0}^{n} w_{j} K(z_{i}, y_{j}) \tilde{f}(y_{j}) = g(z_{i}) \quad (i = 0, 1, 2, ..., m) .$$
 (5.108)

We shall suppose that m=n, until further notice, and these equations then define $\tilde{f}(y_0),\ldots,\tilde{f}(y_n)$ provided $\det \left[K(z_i,y_j)\right] \neq 0$ (in particular, if the functions $K(x,y_j)$ form a Haar set). There is no necessity in these equations to choose $z_i=y_i$, though it may be thought natural to do so. We shall demonstrate by examples that eqns (5.108) do not lead to a generally reliable method of approximating the values of a solution of the integral equation. We shall proceed under the general assumption that K(x,y) and g(x) are continuous.

J. ASSOC. COMPUTING MACHINERY
9,84,97, 1962

A Technique for the Numerical Solution of Certain Integral Equations of the First Kind*

DAVID L. PHILLIPS†

Argonne National Laboratory, Argonne, Illinois

Introduction

The general linear equation may be written as

$$h(x)f(x) + \int_a^b K(x,y)f(y) \ dy = g(x) \qquad (a \le x \le b)$$

where the known functions h(x), K(x, y) and g(x) are assumed to be bounded and usually to be continuous. If $h(x) \equiv 0$ the equation is of first kind; if $h(x) \neq 0$ for $a \leq x \leq b$, the equation is of second kind; if h(x) vanishes somewhere but not identically, the equation is of third kind. If the range of integration is infinite or if the kernel K(x, y) is not bounded, the equation is singular. Here we will consider only nonsingular linear integral equations of the first kind:

$$\int_{a}^{b} K(x,y)f(y) dy = g(x) \qquad (a \le x \le b) \qquad (1)$$

There is extensive literature on equations of the second kind, but literature on linear equations of the first kind is sparse. However, several methods for solving equations of the first kind numerically have been proposed [1-10]. No method has been very successful for arbitrary kernels when the function g(x)is known with only modest accuracy. The reason for this is inherent in the equation itself. Think of the equation as a linear operator, operating on f(y) to produce g(x). This operator does not have a bounded inverse (it may not even have an inverse, but we will assume here that it does) which can be seen as follows. Let f(y) be the solution to (1) and add to it the function $f_m = \sin my$. For any integrable kernel it is known that $g_m \equiv \int_a^b K(x, y) \sin(my) dy \to 0$ as $m \to \infty$. Hence only an infinitesimal change g_m in g causes a finite change f_m in f (i.e. the equation is unstable). Also, one would expect that $g_m \to 0$ as $m \to \infty$ faster for flat smooth kernels than for sharply peaked kernels (indeed if K(x, y)were the δ -function, $K(x, y) = \delta(x - y)$, then $g_m = f_m$ would not approach zero). Hence we conclude that the success in solving equation (1) by any method depends to a large extent on the accuracy of g(x) and the shape of K(x, y).

^{*} Received June, 1961.

[†] Based on work performed under the auspices of the U.S. Atomic Energy Commission.

From equations (13), the matrix method described here merely replaces A by $A + \gamma B$ where B is a certain matrix whose elements depend only on A and γ is an arbitrary non-negative parameter which controls the amount of smoothing. Increasing γ produces greater smoothing. It follows from the second equation in (13) and (5) that e is approximately proportional to γ so that only a very few values of γ need to be used in order to find one giving e the desired magnitude. However, one additional matrix inversion is needed for each new value of γ used. The value of e is determined from the accuracy of the g_i .

Examples -

Example 1. Let the problem be the following:

$$\int_{3}^{3} K(x-\lambda)f(x) \ dx = g(\lambda)$$

where

$$K(z) = 1 + \cos \frac{\pi z}{3},$$
 $|z| \le 3,$
 $= 0,$ $|z| > 3;$
 $g(z) = (6 + \lambda) \left(1 - \frac{1}{2} \cos \frac{\pi \lambda}{3}\right) - \frac{9}{2\pi} \sin \frac{\pi \lambda}{3},$ $|z| \le 6,$
 $|z| \ge 6.$

The solution to this problem is f(x) = K(x). Hence we can easily check the numerical solution against the true solution. Let us first take n = 12 (13 points) and use Simpson's rule for the quadrature formula. The truncation error is about 4. The values of g are rounded off so that the maximum error in $g(\lambda_i)$ is .00005. Table I and Figure 1 give the comparison between the true solution and the numerical solution for several choices of γ . Since the solution is symmetric

TARLE		
	-	

	TABLE I						
$\begin{array}{c c} \lambda & \text{true values} & \gamma = 0 \\ \text{of } i & \end{array}$	$\gamma = .0011$	γ= '.011	$\gamma = .03$	γ = .1	γ = .5 · ·	γ= 1	
0 2.0000 2.943 1 1.5000 1.154 2 .5000 .694 3 .0000 .014 4 .0000 .000 5 .0000 .000 6 .0000 .000	1.404 .388 .077 034 .005	2.082 1.478 .464 002 .031 .029 128	2.017 1.488 .503 016 .019 .038 133	1.948 1.489 .559 014 026 .034 052	1.843 1.476 .651 .016 105 .009	1.796 1.463 .686 .042 120 009	
ave. \(\epsilon_j \) \(\ldots \)	.010	.020	.026	.047	. 105	.140	
$\max \epsilon_j \dots 0$.016	. 035	. 053	.096	. 203	.242	

60

where $g_i = g(x_i)$, $f_j = f(x_j)$ and $A_{ij} = w_j k_{ij}$. It is a simple matter, in principle, to resolve this system (for $m \ge n$) to yield an *n*-point numerical approximation to the solution. In particular, taking n = m yields an 'exact' numerical solution—one that 'models' the discrete data function to within numerical round-off.

At first sight it would appear that the system in equation (5.4) provides a simple, robust numerical inversion technique for resolving equation (5.2). In practice however, the quadrature method runs up against severe difficulties. In particular, when an attempt is made to refine the discretised solution by increasing the number of support points in the source function representation, the conditioning (cf §5.3.2) of the matrix system generally worsens disastrously, leading to a severe magnification of data (and discretisation) errors and eventually to quite meaningless, oscillatory, numerical solutions. For the simplest case of a square matrix system (m = n) the numerical inversion deteriorates as the matrix operator A^{-1} begins to reflect, more and more closely, the unboundedness of the original integral operator \mathcal{K}^{-1} (cf §4.3.1). The problem is generally most severe for strongly smoothing kernels (cf §5.3) but, in any given case, an optimum representation can usually be found by judiciously selecting the spacing of the discretisation mesh (x_i, y_j) in addition to the number of support points (n) in the numerical solution. It is disturbing, none the less, to obtain a numerical formulation of the problem, the solution of which depends critically on what appears to be an essentially arbitrary numerical mesh. It also conflicts with the notions of classical numerical analysis to the extent that numerical solutions are expected to show some degree of convergence as discretisation errors are systematically reduced-provided of course that round-off error is kept at bay.

We return to a more quantitative discussion of the numerical instability

in §5.3.

1' CONCLUSION

NTILIZENDO UN MÉTODO DE INTEGRACION NUMERICA DIRECTA PARA REPRESENTAR NUMERICAMENTE UNA TRANSFORMADA INTEGRAL HAY GENERALMENTE, PROBLEMAS (MUCHOS Y MUY GRAVES) (BA LER) PERO SOBRE TODO CUANDO SE UTILIZA SIMPSON

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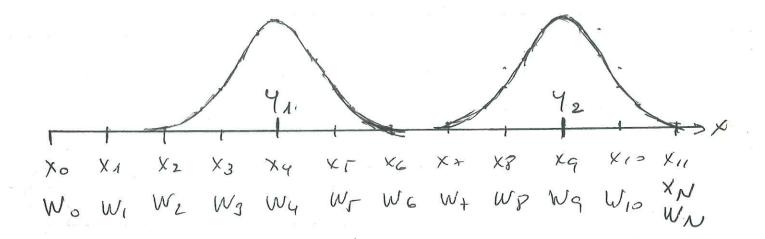
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NUMERICAMENTE

TRAPECIOS

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 $W_1 = W_2 = W_3 - - = W_6 = - W_{N-1} = \Delta$

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$$\frac{N}{\sum_{k=1}^{N} \sqrt{e^{-(k-4)^2}}} = \frac{e^2}{\sqrt{e^2}} + 2\frac{e^2}{\sqrt{e^2}} + 2\frac{e^2}{\sqrt{e$$

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V (x=9)2 = 0,9999999 = 0,9999999

> el results de interne pour live trapacier le gaussiance & ecleute (pres » une consultant

S/MPSON $Wo = \frac{4}{3}J$ $W_{1} = \frac{4}{3}J$ $W_{2} = W_{3} = W_{5} = -\frac{2}{3}J$ $W_{2} = W_{4} = W_{5} = -\frac{4}{3}J$ $\Delta = L$ $Si \qquad Y_{3} \text{ coincide con } x_{4}$ $Le \quad INTEGRAL \quad NUMERICA$

= Wo e + W, e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e + W = e

PEDR QUE CON TRAPECIOS

(pero es casualidad)

(s: A=1/2

Le rune de 1.00034

Si $\frac{1}{1} = \frac{1}{2}$ Le integral rele 0.943567

HUCHISITA DIEERENCIA

L' $1 = \frac{1}{2}$ Le integral rele 0,999965

LUEGO HAT OSCILACIONES

IMPORTANTES EN LOS SALURES

DEPENDIEMDO DE POSÍCION

DE YJ:

"DEL VALOR DE LOS

ALTERNANCIA

2 A GALZAN SALETE
3 A GALZAN SALETE

Y ESTO EN UN CASO MUZ

SIMPLE

PESOS ASOCIADOS

MUCHOS, MUCHISIMOS DE LOS RESULTADOS CATASTROFICOS EN LA INVERSION DE TRANSFORMADAS INTEGRALES PROVIENEN DE LA "DISCRETIZACION" DIRECTA CON PESOS DE INTEGRACION " SIMPSON"

Pero

REPETIDD DE INTEGRACION NUMERICA
UN PROBLEMA MUCHO MÁS DELICADO
A PARECE CUANDO LA PUNCIÓN FIXI
DE PENDE DE UN PARAMETRO X: F(X)
DE TAL FORMA QUE PUEDA
LAMBIAR SU FORMA FUNCIONAL PARA
DIFERENTES VALORES DE ESTE
PARA METRO

ENTONCES PUEDE DOURRIR QUE
PARA CADA VALOR DE LEXISTIRA

UNA FORMA DE INTEGRACION NUMERICA

QUE OPTIMICE LA INTEGRAL

$$\int_{A}^{B} F_{x}(x) dx = \sum_{j=1}^{N} W_{x}(xj) F_{x}(xj)$$

ES DECIR PARA CADA À EL CONJUNTO DE ORDENADAS DISCRETAS Y EL CONJUNTO DE PESOS ASOCIADOS SERA DIFERENTE

Y, A VELES TENEMOS QUE HACER MUCHAS INTEGRALES PARA MUCHAS FS (x) CON EL MISMO CONJUNTO DE ORDENADAS DISCRETAS XXXX Y EL MISMO CONJUNTO DE PESOS ASOCIADOS.

CONCLUSION !!!

EN LAS ECUACIONES INTEGRALES SE PRESENTA, OBVIAMENTE, LA DIFICULTAD ANTERIOR

LA PUNCION FLX) QUE HAY
QUE INTEGRAR, DEPENDE DEL
PARANETRO Y:

K(Y,X) f(x) kx

ASI QUE: - MUCHO CUIDADO - MUCHA EXPERIENCIA O BUSCAR OTROS METODOS.

EN EL CASO EN EL QUE EL
INTERVALO DE INTEGRACIÓN (A,B)

ES FINITO

LA REPRESENTACIÓN FUNCIONAL

DEL INTEGRANDO SE BASA EN

A PROXIMACIONES POUS NO HÍALES

DEL INTEGRANDO

Y UN POLINOMIO ES UN POLINOMIO

AUNQUE SUS COEFICIENTES DEPENDAN

DEL PARÁMETRO Y.

ASI QUE, AL HENOS EN ESTE CASO PUEDE SERVIRNOS ESTE MÉTODO EN EL CASO DE LA CONVOLUCIÓN GOUSSIANA

9(4) = [= (x-4)2 fox dx

PARA CAUA Y TENEMOI UNA GRUSSIANA CENTRADA, PRECISAMBNIE, EN 25E Y

PODREMOS ELEGIR UN CONJUNTO COMUN DE ORDENADAS XE

- DONDE EVALUARAMOS P(YE)
Y EL CONJUNTO DE PESOS

CORRESPONDIENTE WE

PERO ESOS PESOS WE

ASOCIADOS AL CONJUNTO ME

TIENEN QUE SER TALES

Je-(x-4)e dx = = = (xx-4) Wx

SEA IGUAL PARA TODO Y LO QUE NO OCURRÍA ANTES

POR EJEMPLO

TRANSFORMADA DE LAPLACE

UN SISTEMA DE ORDENADAS DISCRETAS XXXX TALES QUE

9(4) = Z W(K) @ XKY f(Xe)

SEA OPTIMA

DEPENDERA EXTRAORDINARIAMENTE DE Y

Para cale 9 tendreurs me compute option de valore XIC 9. Wa Pers tempo que violver el pollema con los valores de XE 9 de We dos veisues para toda los 9.